

1 **Poleward Transport of African Dust to the Iberian Peninsula Organized by a**
2 **Barrier Jet and Hydraulic Jumps: Observations and High-Resolution**
3 **Simulation Analyses**

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20 **Highlights:**

- 21 • Saharan dust episode with a strong impact poleward over the Iberian Peninsula.
- 22 • High-resolution WRF-Chem simulations of a severe African dust storm.
- 23 • Multi-scale terrain-induced circulations are instrumental for dust ablation.
- 24 • Barrier jet and hydraulic jumps organize dust storms.
- 25 • Hydraulic jumps are linked to a mesoscale gravity wave.

26 **Abstract**

27 Most air quality stations in Spain exceeded the European Union's daily PM₁₀ limit due to the
28 February 2016 Saharan dust outbreak, which resulted from two successive dust storms in
29 Northwest Africa. This study identifies the meso- β /meso- γ -scale dynamical processes
30 responsible for developing these dust storms using observations and high-resolution Weather
31 Research and Forecasting model coupled with Chemistry simulations. The results revealed that
32 the first dust storm was associated with a strong barrier jet (BJ). The BJ formed on the
33 southeastern foothills of the Saharan Atlas Mountains (SAM) when an easterly-northeasterly
34 low-level Mediterranean flow was blocked by a stably stratified layer close to the SAM. The BJ
35 intensified just after sunrise on 20 February and the associated near-surface peak winds
36 organized the first dust storm. The second dust storm was linked to a mesoscale gravity wave
37 (MGW) and hydraulic jumps. A long-lived westward propagating MGW was triggered by a
38 downslope flow interacting with the stable layer near the northeastern edge of the Tinrhert
39 Plateau in eastern Algeria. When this MGW crossed the Tademaït Plateau, hydraulic jumps
40 formed on its lee side. The strong winds accompanying these hydraulic jumps formed the second
41 dust storm on 21 February. The lifted dust extended over a depth of 2-3 km in the growing
42 daytime boundary layer and was advected poleward by the southerly/southeasterly mid-
43 tropospheric winds. Our results underline the importance of resolving terrain-induced mesoscale
44 processes to understand dust storm dynamics, which are difficult to represent in coarse-
45 resolution numerical models.

46 **Keywords:** Saharan dust storm, terrain-induced circulations, barrier jet, hydraulic jumps, WRF-
47 Chem

48 **1 Introduction**

49 Terrain-induced meso- β /meso- γ -scale meteorological features are known to organize
50 severe dust storms over North Africa. Known mechanisms include downslope winds associated
51 with hydraulic jumps (e.g., Gläser et al., 2012; Pokharel et al., 2017), convectively generated
52 density currents linked through orographic forcing (Knippertz et al., 2007; Roberts & Knippertz,
53 2014), and density current-like cold fronts and undular bores (e.g., Dhital et al., 2020). A
54 regional model simulation by Gläser et al. (2012) identified strong near-surface winds
55 accompanying a hydraulic jump on the southern slope of the SAM that emitted dust during the
56 early stage of the March 2004 Saharan dust outbreak. Similarly, Pokharel et al. (2017) showed
57 that downslope windstorms and hydraulic jumps on the southern slope of the SAM and
58 southwestern slopes of the Tibesti Mountains in Chad occurred as a small-scale and isolated
59 precursor wind event before the formation of a large-scale dust storm. Dhital et al. (2020) found
60 that a density current-like cold front induced an undular bore that organized the 10-13 November
61 2017 strong dust episode on the southern flank of the SAM, which ultimately impacted the Cape
62 Verde Islands. The study by Roberts and Knippertz (2014) on the June 2010 intense Haboob near
63 the Hoggar and Air Mountains found that a terrain forcing mechanism favored the formation of
64 convective cold pools and associated strong winds critical for dust emission. Knippertz et al.
65 (2007) suggested that instability paired with a relatively high moisture content over the Saharan
66 Atlas triggered deep moist convection and an associated dust-emitting density current.

67 Another phenomenon of potential interest in complex terrain is the barrier jet (BJ), which
68 represents a class of strong low-level mesoscale winds that develop adjacent to steep terrain
69 (e.g., Parish 1982). A BJ forms when the low-level upstream flow is blocked (e.g., Loescher et

70 al., 2006). This requires a low-level stable stratification that enhances the pressure gradient force
71 parallel to the windward side of a mountain barrier. In response, the flow is accelerated and
72 redirected parallel to the mountain barrier. Given the necessary stable stratification, BJs are
73 frequent during the cool season (e.g., Colle et al., 2006; Neiman et al., 2010). They have, for
74 instance, been observed over the Appalachian Mountains (e.g., Bell and Bosart 1988), in the
75 Colorado Rockies (e.g., Cox et al., 2005), the western slope of the Sierra Nevada Mountains
76 (e.g., Parish 1982; Neiman et al., 2010), the Pacific Northwest (e.g., Braun et al. 1997), the
77 Alaskan coast (e.g., Loescher et al., 2006, Colle et al., 2006;), the Gulf of Mexico (Luna-Niño &
78 Cavazos, 2018), and Taiwan (Li and Chen, 1998).

79 Another ubiquitous feature in complex terrain is the hydraulic jump which forms in
80 conjunction with downslope winds on the lee side of a mountain barrier when an upstream flow
81 has enough kinetic energy to overcome the potential energy associated with a mountain barrier.
82 Besides North Africa, downslope winds and hydraulic jumps have been observed over the
83 eastern Sierra Nevada and Owens Valley, California (Lin, Y., 2007), the front range of the
84 Rocky Mountains (Karyampudi et al., 1995), northern Australia (e.g., Clark, 1972), and the
85 Middle East (e.g., Pokharel et al., 2017).

86 The dimensionless Froude number, $Fr = U/(Nh)$, has been frequently used to characterize
87 the upstream flow regime (e.g., Kaplan et al., 2012; Pokharel et al., 2017), where U is the
88 ambient wind speed perpendicular to the mountain barrier, N is the Brunt-Väisälä frequency, and
89 h is the barrier height. An upstream subcritical flow ($Fr < 1$) favors BJ formation, while
90 supercritical flow ($Fr > 1$) favors hydraulic jumps and associated downslope wind formation. In
91 the hydraulic jump region, the flow changes from a supercritical regime to a subcritical regime.

92 Severe weather related to BJs and hydraulic jumps have been documented in previous
93 studies. For example, Neiman et al. (2013) showed that an interaction between the Sierra barrier
94 jet, i.e., the BJ on the western side of the Sierra Nevada Mountain range, and an atmospheric
95 river resulted in heavy precipitation and flooding events in the Central Valley of California. Ke
96 et al. (2019) identified that the interaction between a BJ and a cold pool outflow associated with
97 mesoscale convection results in strong convection and heavy precipitation over northern Taiwan
98 during the Mei-Yu season, i.e, mid-May to mid-June. Similarly, outside North Africa, the
99 importance of hydraulic jumps and an undular bore in a severe weather outbreak has also been
100 highlighted. Karyampudi et al. (1995), in their study of the April 1986 severe weather outbreak
101 over eastern Colorado and western Nebraska, identified the formation of a downslope windstorm
102 associated with a hydraulic jump and undular bore to the east of the Rockies when a density-
103 current like cold front spilled onto the lee side of the mountains. The interaction between a
104 downstream propagating undular bore, lee cyclone, a dryline, and a warm front triggered deep
105 convection as well as intense blowing dust in this case.

106 A severe African dust outbreak over the IP occurred during 20-21 February 2016. Dust
107 plumes were swept by zonal flows eastwards on subsequent days and eventually impacted
108 Germany (https://www.dlr.de/eoc/en/desktopdefault.aspx/tabid-11128/19488_read-45631/). A detailed
109 ground-based observational analysis on this outbreak by Titos et al. (2017) found that ~90 % of
110 the air quality stations in Spain exceeded the European Union's daily PM_{10} limit of $50 \mu g \cdot m^{-3}$.
111 Gama et al. (2019) also highlighted the severity of this dust outbreak over Portugal on air quality
112 based on observational and modeling studies. However, this study did not focus on the detailed
113 analysis of the dust emission processes. In a companion paper (Orza et al., 2020; hereafter
114 referred to as Part I), the synoptic analysis of this dust episode was presented in detail. During

115 this episode, two distinct dust plumes reached the IP in succession-the first of two dust plumes
116 crossed the SAM during midday on 20 February and was advected towards the Western IP. The
117 second one followed in the afternoon of 21 February and was advected towards the Eastern IP.
118 Part I highlighted a double Rossby Wave Break (RWB) in the polar jet stream (PJ) linked
119 through nonlinear wave reflection as an upper-level synoptic precursor flow that favored this
120 dust storm formation and subsequent poleward transport of dust to the IP.

121 In this paper, we extend our analyses to meso- β /meso- γ -scale meteorological processes
122 and highlight the importance of a BJ and hydraulic jumps in organizing this dust outbreak to
123 Europe. We find that the BJ is another important terrain-induced mechanism for severe dust
124 outbreaks. The BJ results in strong near-surface winds capable of emitting dust from potential
125 source regions on the windward side of a mountain barrier. For perhaps the first time for the
126 entire of North Africa, this study documents a BJ and assesses the role in organizing a strong
127 dust storm. In addition to the BJ, this study also provides evidence of a subsequently occurring
128 severe dust storm organized by hydraulic jumps in association with a long-lived mesoscale
129 gravity wave (MGW). MGWs are gravity waves, with a wavelength longer than 50 km and
130 periods exceeding 1 hr (Ruppert and Bosart, 2014), which have been known for their role in
131 different severe weather events (e.g., Uccellini and Koch, 1987; Kaplan et al., 1997), but not for
132 North African dust storms.

133 The remainder of this paper is organized as follows. Section 2 provides an overview of
134 our data, methods, and model configuration. The synoptic setup for BJ, mesoscale wave
135 dynamics, and observational analysis of the spatiotemporal evolution of the dust storm is
136 presented in Section 3. Section 4 outlines the detailed dynamics of the BJ and hydraulic jumps,
137 highlighting their importance in this dust storm occurrence. The summary and conclusions of the
138 present study are presented in Section 5.

139 **2 Data and methods**

140 2.1 Observations

141 Wind speed, wind direction, and visibility data for the observational network distributed
142 over the study area (Figure 2a), for 20-21 February 2016, were obtained from the Meteorological
143 Terminal Aviation Routine Weather Report
144 (<https://mesonet.agron.iastate.edu/request/download.phtml>) and provided the time and location
145 of the reduced visibility associated with this dust storm. The combined Dark Target Deep Blue
146 aerosol optical depth (τ) from the Moderate Resolution Imaging Spectroradiometer aboard the
147 Aqua satellite (<https://ladsweb.modaps.eosdis.nasa.gov/>) and sounding data from the University
148 of Wyoming (<http://weather.uwyo.edu/upperair/sounding.html>) provided the spatial extent of the
149 dust evolution and vertical profiles for the model validation. The large-scale flow features that
150 create a favorable environment for BJ formation were described using the European Center for
151 Medium-Range Weather Forecast (ECMWF) ERA-Interim reanalysis dataset (Dee et al., 2011).
152 We used the charts of potential vorticity (PV) and horizontal wind at the 330 K isentropic surface
153 and 850 hPa temperature and wind.

154 2.2 Model configuration

155 Meso- β -/meso- γ -scale meteorological processes are analyzed in high-resolution
156 simulations with a one-way nesting approach using the WRF-Chem model version 3.9 (Grell et

157 al., 2005). The outer, middle, and inner domain's horizontal resolutions are 18, 6, and 2 km,
158 respectively, and covered the area as shown in Figure 2a. All the model simulations have 40
159 vertical levels, with higher resolution in the lower atmosphere and the vertical boundary
160 condition at 50 hPa. The parent domain is initialized using the ERA-Interim reanalysis data,
161 while the inner nested domains are initialized by the simulation results from the outer domain.
162 The outer, middle, and inner domains are initialized at 00 UTC, 06 UTC, and 12 UTC on 19
163 February, respectively, and all end at 06 UTC on 21 February.

164 In this study, the convective parameterization is employed in the 18 km simulation only
165 using the Betts-Miller-Janjic scheme (Janjic, 1994) and turned off in the high-resolution domains
166 (6 and 2 km) thus explicitly resolving moist convection on the model grid. Our model
167 configurations are otherwise identical across the different resolutions. The model configuration
168 uses the double-moment bulk microphysical parameterization (Thompson et al., 2008), the
169 Mellor-Yamada-Janjic (MYJ) PBL scheme (Mellor & Yamada, 1974; Janjic, 2002), Noah Land
170 Surface Model (Chen & Dudhia, 2001; Ek et al., 2003), the Dudhia shortwave radiation scheme
171 (Dudhia, 1989), and the RRTM for longwave radiation (Mlawer et al., 1997). These
172 parameterizations have been successfully used to simulate several terrain-induced mesoscale
173 meteorological features organizing strong dust storms over North Africa (e.g., Pokharel et al.,
174 2017; Dhital et al., 2020).

175 The WRF-Chem simulation is performed in dust-only mode following the Georgia
176 Tech/Goddard Chemistry Aerosol Radiation and Transport (GOCART) dust scheme (Ginoux et
177 al., 2001). The GOCART scheme includes five dust bins having effective radii of 0.73, 1.4, 2.4,
178 4.5, and 8 μm . The model outputs dust τ at 550 nm using the corresponding columnar mass load
179 and the extinction efficiencies at 550 nm and dust concentration in each size bin. The simulated
180 dust emission flux (F_p) in the GOCART dust scheme is a function of wind speed, erodibility, and
181 surface wetness and is calculated using equation (1).

$$182 F_p = CS s_p u_{10m}^2 (u_{10m} - u_t) \quad \text{if } u_{10m} > u_t \quad (1)$$

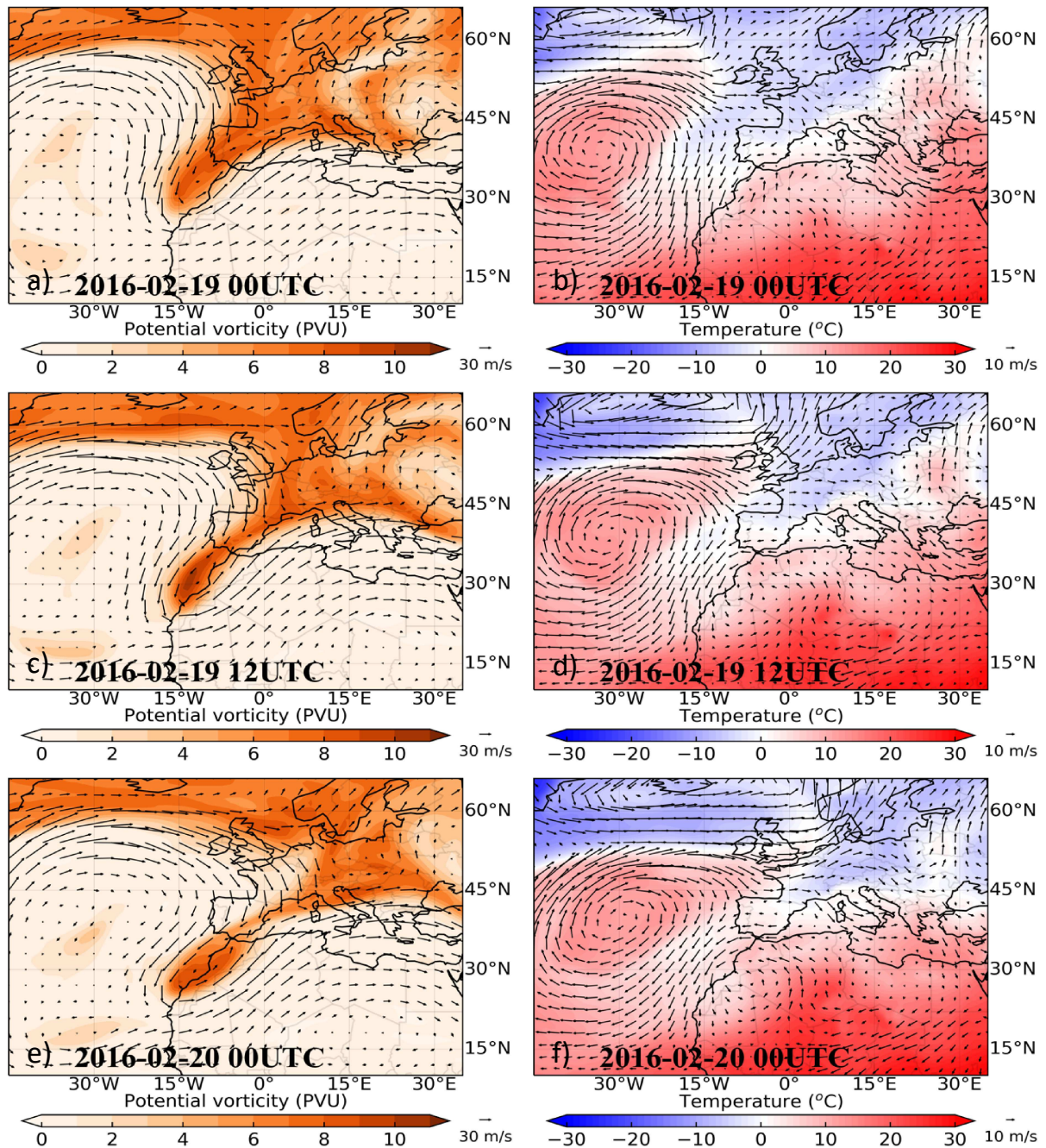
183 where the constant $C = 1 \mu\text{g}\cdot\text{m}^{-5}\cdot\text{s}^2$, S is source function, s_p is the mass fraction of size group p
184 of dust emission, u_{10m} is the wind at 10m height, and u_t is the threshold wind velocity for the
185 effects of wind erosion. The source function (S) is a dimensionless quantity, which depends upon
186 the soil properties. The meteorological interpretations are made based on 2 km simulation
187 outputs unless otherwise stated.

188 **3 Observed downscale evolution from RWB to mesoscale dust plumes**

189 **3.1 Synoptic setup for barrier jet and mesoscale wave dynamics**

190 Figure 1 depicts the 330K winds and isentropic PV as well as the 850 hPa temperatures
191 and wind vectors during the precursor period from 00 UTC 19-20 February. In this period, the
192 favorable thermodynamic profile for the BJ develops above the Atlas Mountains. The 330K
193 fields indicate an anticyclonic RWB in the PJ above the western European Coast extending
194 offshore and south southwestward. This feature subsequently propagates southeastward towards
195 the northwestern African Coast during this 24-hour period (Figures 1a, c, and e). The RWB and
196 PV reversal is the result of a tongue of PV wrapping anticyclonically into northwestern Africa
197 and is in proximity to the subtropical jet stream (STJ) exhibiting substantial southwesterly flow.
198 This southwesterly flow creates a strong confluence and shearing deformation zone from the

199 tropopause to the lower troposphere over northwestern Africa. At the same time, one sees the
200 intensifying low-level baroclinic zone at 850 hPa and that surface's confluent wind vectors
201 centered near the SAM (Figures 1b, 1d, and 1f). This baroclinic zone is the result of the
202 frontogenesis in the aforementioned confluence/shear zone in the low-middle troposphere as: 1)
203 polar air over northern Europe is advected south-southwestward within the RWB in proximity to
204 2) hot continental tropical (CT) air above the Sahara Desert being advected towards and over the
205 Atlas Mountains. This deep frontogenetical zone, where north-northeasterly low-level flow
206 caused by the PJ RWB and the low-level south-southwesterly return branch jet in the STJ
207 converge, is critical for setting up the BJ. The favorable conditions for the BJ are established as
208 the very warm CT air under the STJ overruns cooler Mediterranean air emanating from the
209 north-northeast under the PJ thus resulting in the statically stable and veering flow described
210 earlier above the Atlas Mountains which is so conducive to blocking and BJ formation. The large
211 magnitude static stability results from the intensifying frontogenetical circulation accompanying
212 the overrunning of CT air on the equatorward slope of the Atlas Mountains in Algeria and
213 Morocco. The hot, dry CT air plume originating as part of the low-level return branch in the 700-
214 850 hPa layer under the STJ from the western Sahara.



215

216 Figure 1. PV and wind vector at the 330 K isentropic surface (left). Temperature and wind vector
 217 at the 850 hPa level (right). (a, b) at 00 UTC, (c, d) at 12 UTC on 19 February, and (e, f) at 00
 218 UTC on 20 February.

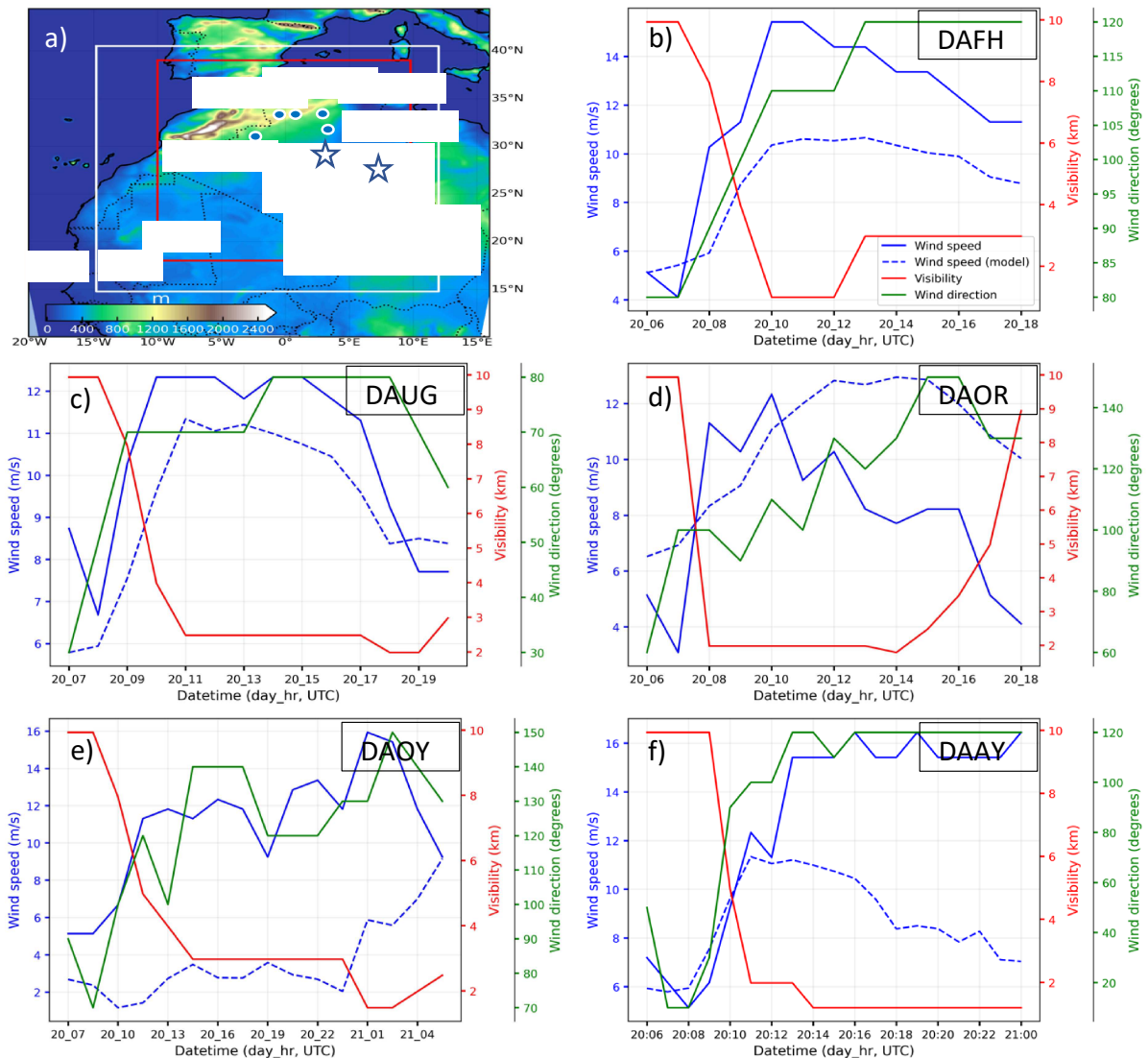
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3.2 Spatiotemporal evolution of the observed dust outbreak

220

Our observational analyses suggest that this dust episode was associated with two
 221 successive dust storms initiated on the southeastern side of the SAM. The first dust storm started
 222 ~09 UTC on 20 February and second storm began ~01 UTC on 21 February. The lifted dust then

223 ultimately was advected poleward towards the IP (see Figure 3 in Part I). Figure 2 illustrates the
224 observed wind speed, wind direction and visibility for 20-21 February 2016 at different stations
225 in Algeria. The initial signal of the dust storm was recorded at two stations [Tilrempt (DAFH,
226 32.93°N, 3.31°E) in Laghouat Province and Béchar (DAOR, 31.62°N, 2.23°W) in Béchar
227 Province] at 08 UTC on 20 February (Figures 2b and 2d). The DAFH station recorded an
228 easterly wind speed of 10.29 m.s⁻¹ and visibility of 4.97 km, while the DAOR station recorded an
229 east-southeasterly wind speed of 11.3 m.s⁻¹ and visibility of 1.24 km. At 09 UTC, the DAFH
230 station recorded an east-southeasterly wind speed of 11.3 m.s⁻¹ and visibility of 2.49 km while
231 the Ghardaia station (DAUG, 32.38°N, 3.79°E) in Ghardaïa Province recorded a northeasterly
232 wind speed of 14.4 m.s⁻¹ and visibility of 4.97 km (Figures 2b and 2c). The visibility was
233 unchanged at the DAOR station with an easterly wind speed of 10.29 m.s⁻¹. At 10 UTC, the wind
234 speed at three stations (DAFH, DAUG, and DAOR) further increased and the visibility remained
235 <3 km. At 11 UTC, a significant reduction in visibility was observed at the Mecheria station
236 (DAAY, 33.58°N, 0.28°W) in Naâma Province in the Atlas Mountains which recorded an
237 easterly/southeasterly wind speed of 12.34 m.s⁻¹ and visibility of 1.98 km (Figure 2f), pointing to
238 the occurrence of desert-dust aerosols. Concurrently, the El Bayadh station (DAOY, 33.67°N,
239 1°E) in El Bayadh Province reported a dust storm associated with a southeasterly wind speed of
240 11 m.s⁻¹ and visibility of 3.11 km (Figure 2e). The visibility at the DAFH, DAUG, and DAOR
241 stations remained <5 km for the 12-18 UTC period. A continuous reduction in visibility was
242 observed at the DAOY (<3km) and DAAY (<2.5km) stations with a southeasterly wind speed of
243 >9 m.s⁻¹ until 15 UTC on 21 February. Additionally, starting 03 UTC on 21 February, the
244 visibility at the DAOR station again started to decrease and remained <3 km till 14 UTC.



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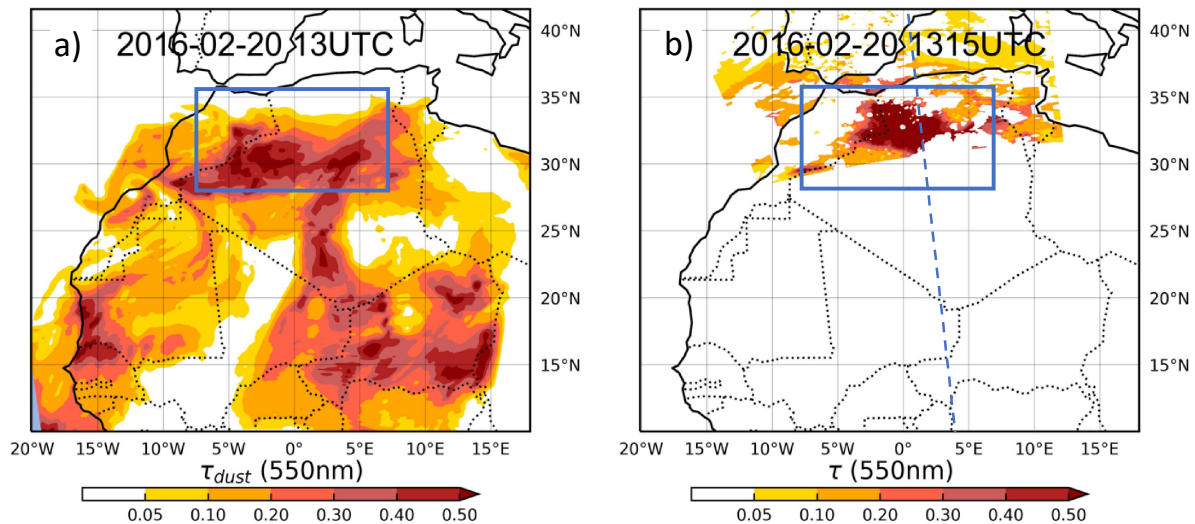
246 Figure 2: (a) WRF-Chem simulation domain, where d1, d2, and d3 represent the domains of 18,
 247 6, and 2 km horizontal resolution, respectively. Time series of wind speed, visibility, and wind
 248 direction are shown for the stations in (b) DAFH, (c) DAUG, (d) DAOR, (e) DAOY, and (f)
 249 DAAY. The dots indicate the locations of the available hourly meteorological stations and stars
 250 mark the center of plateaus mentioned in the text. TP is an abbreviation for the Tadmait Plateau.

251 **4 High-resolution simulation results**

252 4.1 Model validation with observations

253 The simulated dust τ qualitatively reproduces the observed dust plume. A thick dust
 254 plume, as indicated by the higher values of simulated $\tau > 0.5$, was present over the Atlas
 255 Mountains (blue box in Figure 3a), consistent with the reduced visibility at the DAOR station
 256 ~ 13 UTC on 20 February. After 15 min, the MODIS-Aqua overpass at 1315 UTC also captured
 257 the $\tau > 0.5$ over the same region (blue box in Figure 3b). The observed τ was higher over the

258 Atlas Mountains range outside of the marked region and suggests an underestimation of the
 259 simulated dust τ over the ocean regions to the North. The observations, however, include all
 260 aerosol species. An underestimation of dust τ in the model compared to total τ in the
 261 observations is, therefore, to be expected. For instance, the simulated dust τ does not account for
 262 the anthropogenic aerosols and sea-salt aerosols. The τ pattern from MODIS is due to clouds (see
 263 Figure 3 in Part I).

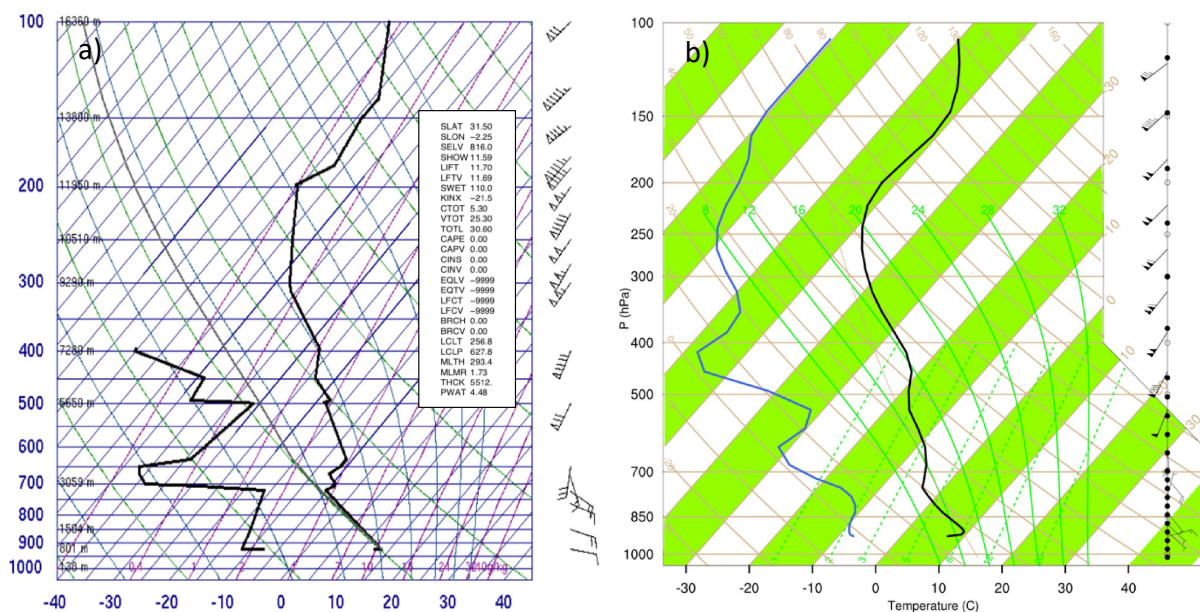


264
 265 Figure 3: Aerosol optical depth. Shown are the (a) WRF-Chem simulated (18 km) dust τ at 13
 266 UTC on 20 February 2016 and the (b) Dark Target Deep Blue combined τ at 550 nm from the
 267 MODIS aboard Aqua at 1315 UTC 20 February 2016. The dashed blue line represents the
 268 MODIS Aqua overpass. The blue boxes mark the τ over the Atlas Mountains discussed in the
 269 text.

270 The comparison between observed and simulated 10m wind speed at three stations,
 271 namely DAFH, DAUG, and DAOR, shows that the simulated wind speed generally followed the
 272 observed wind pattern (Figures 1b-d). At 09-10 UTC February 20, when BJ formed, and
 273 visibility started to decrease, the observed wind speed was $\sim 10-15 \text{ m.s}^{-1}$, and the simulated wind
 274 speed was $\sim 8-11 \text{ m.s}^{-1}$, suggesting an underestimation of simulated wind speed. During 10-16
 275 UTC, the simulated wind speed at the DAFH and DAUG was between $10-11.5 \text{ m.s}^{-1}$, while the
 276 observed wind speed ranged between $11.5-15 \text{ m.s}^{-1}$ (Figure 1b-c). However, at the DAOR station,
 277 the simulated wind speed remained between $11-13 \text{ m.s}^{-1}$ and the observed wind speed ranged
 278 between $8-12 \text{ m.s}^{-1}$ (Figure 1d), suggesting an overestimation of the simulated wind speed. Over
 279 the Saharan Atlas region, at the DAAY station, the simulated wind speed closely matches with
 280 the observation during 09-12 UTC (Figure 1e). At a later time, the simulated wind speed does not
 281 follow the observed wind speed pattern at the same station and at the DAOY station (Figure 1e-f).
 282 The models' inability to generate a similar wind evolution at these two stations is probably
 283 because the model could not capture small-scale features over the Saharan Atlas Mountains
 284 complex topography. To summarize, the evolution of the simulated wind speed on the southern
 285 flank of the Atlas Mountain, where the dust storm was formed, generally followed the
 286 observation with sufficient magnitude to emit dust from the source region providing confidence
 287 in our simulation to assess the dynamics of dust storm formation.

288 Additionally, the observed and simulated temperature and wind profiles at the Béchar
 289 radiosonde station in Algeria at 00 UTC on 20 February are in close agreement (Figures 4a-b).

290 One can notice the presence of (1) a near surface stable layer, (2) vertically veering flow with a
 291 secondary elevated inversion layer, and (3) sharp changes in winds in both soundings. However,
 292 differences also exist between simulated and observed soundings. Near the surface, the simulated
 293 air temperature is $\sim 8^{\circ}\text{C}$, and the observed temperature is $\sim 12^{\circ}\text{C}$. At ~ 900 hPa, the simulated and
 294 observed air temperatures are $\sim 10^{\circ}\text{C}$ and $\sim 11^{\circ}\text{C}$, respectively. Also, ~ 650 hPa level, the observed
 295 and simulated dew point depressions are $\sim 35^{\circ}\text{C}$ and $\sim 23^{\circ}\text{C}$, respectively, which indicates a more
 296 dry layer in the observation than in the simulation. Still, both soundings show a similar pattern of
 297 the near-surface stable layer and the vertical wind profiles necessary to make a favorable
 298 environment for BJ formation. The good agreement with the observations indicates that the
 299 complex vertical atmospheric structure, important for the mesoscale process analysis, is
 300 reproduced by our simulation.



301
 302 Figure 4: Skew T-Log P diagram at the Béchar station at 00 UTC on 20 February 2016 from the
 303 (a) observed sounding and (b) WRF-Chem simulation.

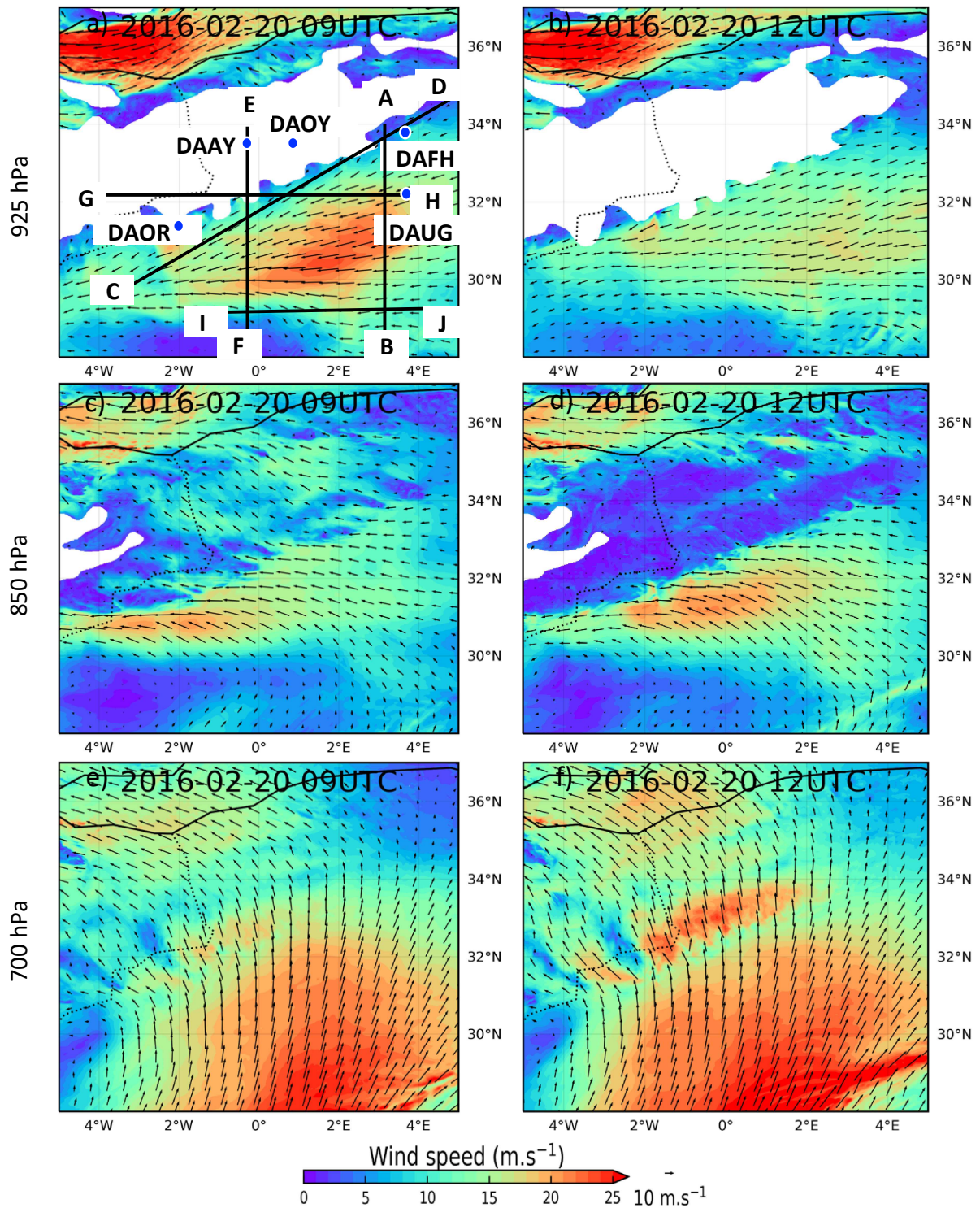
304 4.2 Simulated dust evolution

305 The evolution of the dust outbreak is analyzed using the horizontal winds and dust load at
 306 different atmospheric levels. At 09 UTC on 20 February, winds exceeding $25 \text{ m}\cdot\text{s}^{-1}$ were
 307 observed over the $\sim 30\text{-}32^{\circ}\text{N}$, $0\text{-}2^{\circ}\text{E}$ region (Figure 5a). On the southeastern foothills of the SAM,
 308 the wind speed was $> 15 \text{ m}\cdot\text{s}^{-1}$. These strong winds emitted dust aerosols (Figure 6a and 7a) and
 309 formed the first dust storm. Concurrently, dust loading on the southeastern foothills of the SAM
 310 was $\sim 0.5 \text{ g}\cdot\text{m}^{-2}$.

311 Different wind directions are observed at different heights, suggesting the advection of
 312 lifted dust towards different height-dependent directions. At the 850 hPa level, the southeasterly
 313 flow turned to easterly when it reached the windward side of the Atlas range and the flow
 314 became stronger near Béchar Province (Figure 5c). However, at the 700 hPa level, which is
 315 above the mountain top, the southerly/southeasterly flow prevailed in the region (Figure 5e).

316 Here, the emitted dust was mostly advected downwind of the Béchar station because of the
317 predominantly low-level northeasterly flow (Figures 5a and 5c).

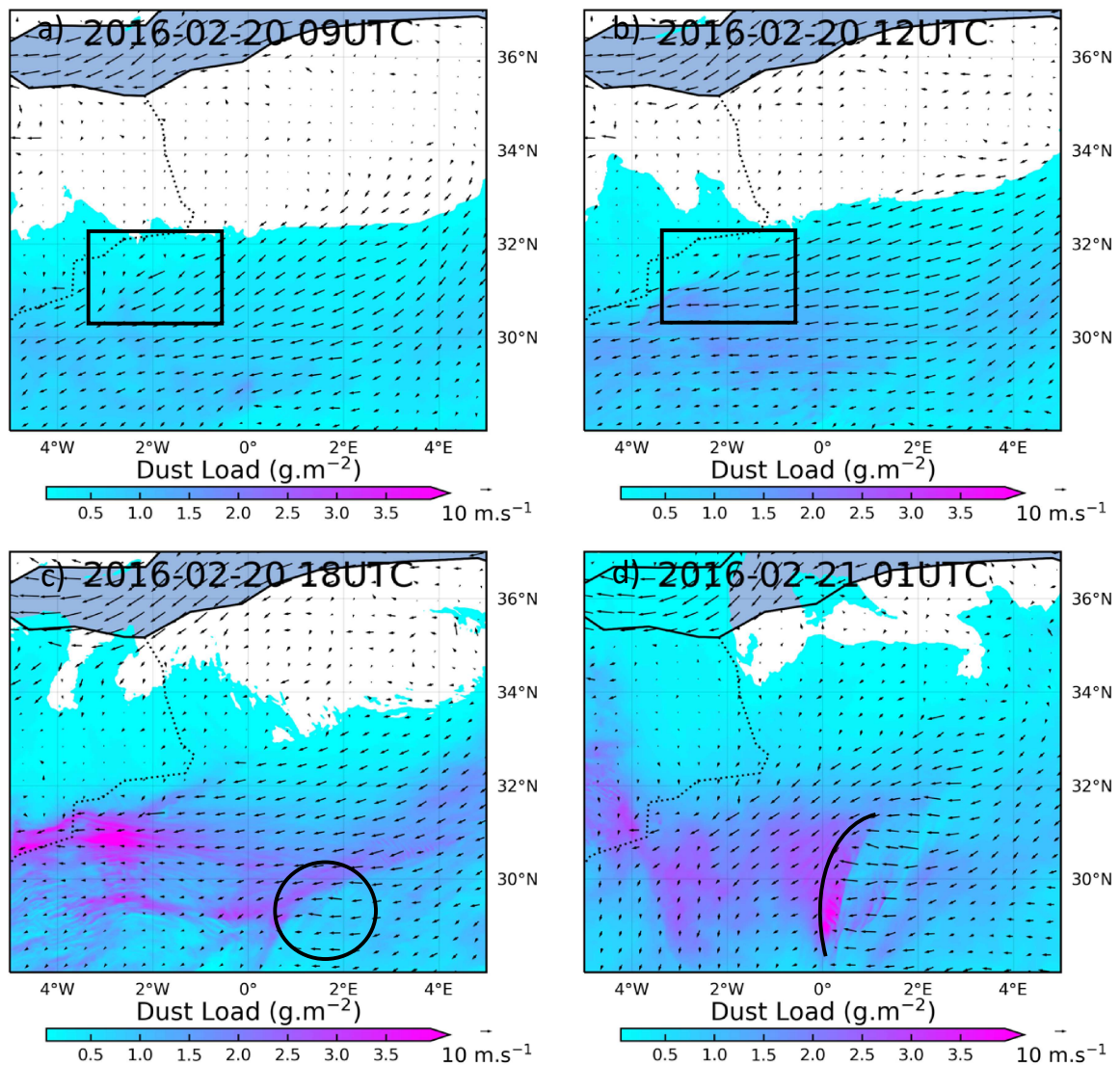
318 The freshly emitted dust did not reach high altitudes initially such that no significant
319 poleward advection by mid-tropospheric winds occurred at this stage (Figures 5e, 6a, and 7a),
320 due to the near-surface stable layer in the morning. At 12 UTC, noticeable features were
321 observed in the wind field at different atmospheric levels. The strength of the near-surface wind
322 decreased (Figure 5b), but the wind speed at higher atmospheric levels and the dust load
323 increased. This indicates upward mixing of the dust by midday when the daytime heating had
324 eroded the surface inversion and dry convection mixed the dust upward. Both the 850 and 700
325 hPa winds strengthened over Béchar Province. However, the wind fields were different: east-
326 south-easterly at 850 hPa and southerly at 700 hPa (Figures 5d and 5e). A significant amount of
327 dust started to advect poleward with an increasing dust load over Béchar Province (see
328 rectangular box in Figure 6b). The dust loading over the top of the SAM was $>0.5 \text{ g.m}^{-2}$
329 consistent with the strengthening of 700 hPa southerly/southeasterly wind (Figures 5f and 6b).
330 These mid-tropospheric winds advected the dust poleward.



331
 332 Figure 5: WRF-Chem simulated horizontal wind vector and speed (fill in m.s^{-1}) at (a, b) 925 hPa,
 333 (c, d) 850 hPa, and (e, f) 700 hPa and valid for 09 UTC (left) and 12 UTC (right) on 20 February
 334 2016. The solid black lines mark cross-sections used in later analyses: A-B (28°N , 3°E to 34°N ,
 335 3°E), C-D (30°N , 3.5°W to 34.5°N , 5°E), E-F (28°N , 0.28°W to 34°N , 0.28°W), G-H (32.38°N ,

336 4°W to 32.38°N, 4°E), I-J (29°N, 1°W to 29°N, 7°E). The white area represents elevation above
337 the pressure levels.

338 At the beginning of the second dust storm, the dust load started to increase on the western
339 side of the TP. At ~18 UTC on 20 February, the 10m wind suddenly began to increase on the
340 west side of the TP, which emitted dust aerosol (black circle in Figure 6c). Synchronously, dust
341 loading again started to increase near Béchar Province, where the dust load reached $>3.5 \text{ g.m}^{-2}$.
342 The 10m wind further strengthened on the lee side of the TP by 01 UTC on 21 February resulting
343 in the second strong dust storm (solid black line in Figure 6d). Afterward, a dust front, or first-
344 order discontinuity in the concentration of dust, moved northwestward towards the Béchar
345 Province. Subsequently, the lifted dust mixed into the growing daytime planetary boundary layer
346 (PBL) after sunrise and was then followed by the southerly/southeasterly wind which advected
347 the dust poleward to the IP. At ~01 UTC on 21 February, the dust plume from the first dust
348 storm had already reached the southern IP at that time (Figure 6d). The second dust plume
349 crossed over the Atlas range at ~15 UTC on 21 February (Figure 3 in Part I). We investigate the
350 mesoscale processes that organized both dust plumes next.



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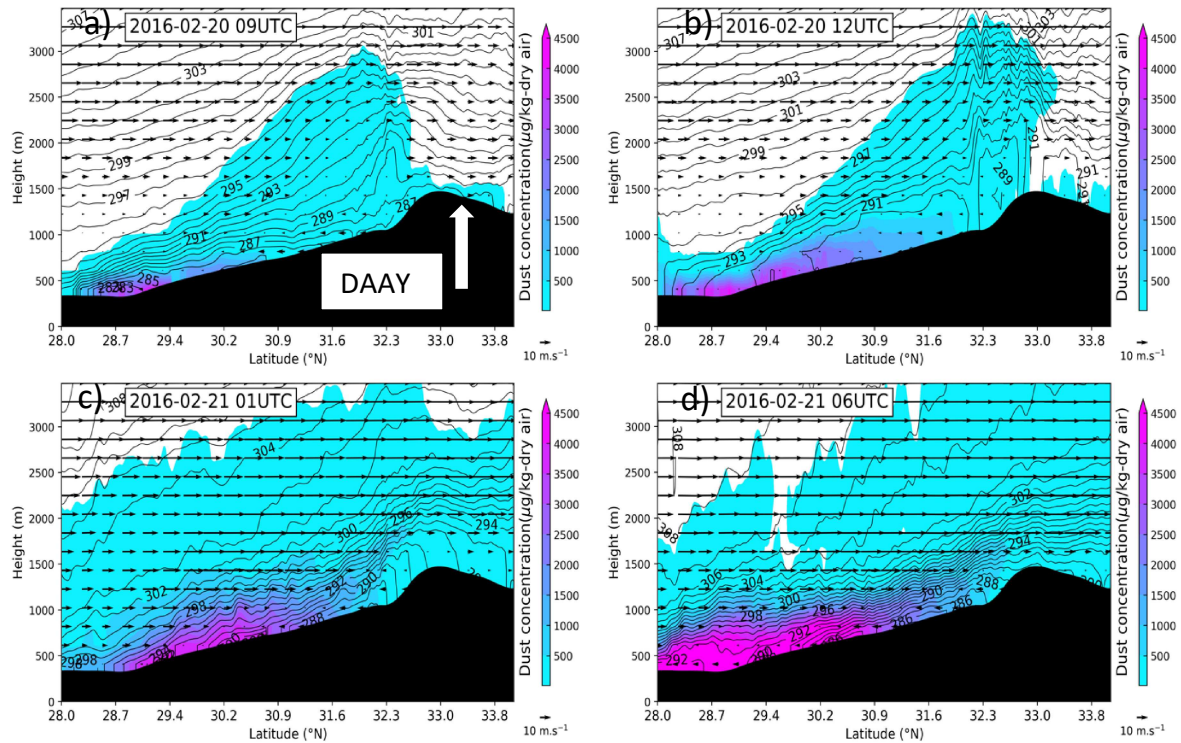
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Figure 6: WRF-chem simulated 10m wind vector (m.s^{-1}) and dust load (fill in g.m^{-2}) at (a) 09 UTC, (b) 12 UTC, (c) 18 UTC on 20 and (d) 01 UTC on 21 February 2016. The black boxes in Figures c and d represent the Bechar Province region. The black circle in Figure c marks the region of increasing 10m wind. The black line in Figure d marks the second dust front.



356
 357 Figure 7: North-South transect at 0.28°W . Shown are WRF-Chem results for the vertical cross-
 358 section of potential temperature (K), wind vector (v and w , $\text{m}\cdot\text{s}^{-1}$), and dust concentration (fill in
 359 $\mu\text{g}\cdot\text{kg}^{-1}$) at the line E-F at (a) 09 UTC and (b) 12 UTC on 20 February 2016, and (c) 01 UTC and
 360 (d) 06 UTC on 21 February 2016. Location of transect is marked in Fig. 5a. White arrow mark
 361 the locations of stations

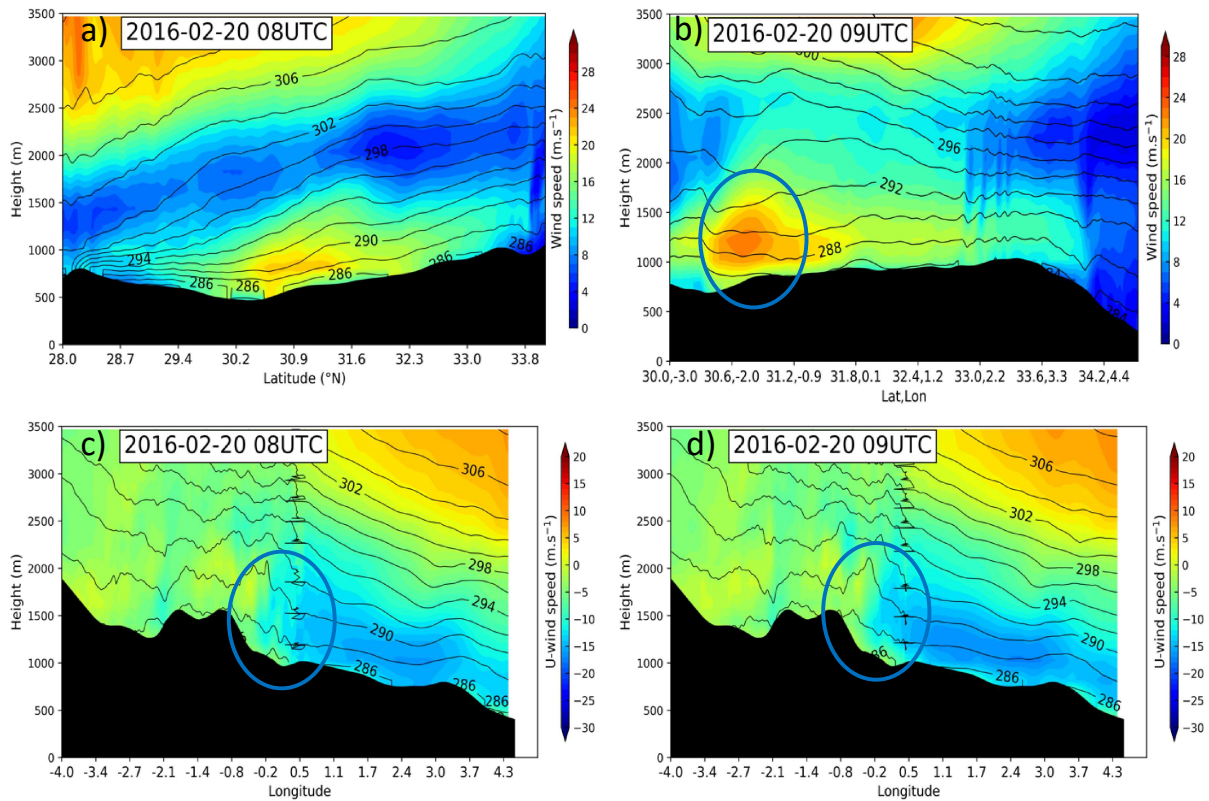
362 4.3 Meso- β /meso- γ -scale dynamics

363 The simulation results showed that the first dust storm was associated with BJ formation
 364 on the southeastern foothills of the SAM. The second dust storm was linked to the evolution of a
 365 long-lived westward propagating MGW near the northern edge of the Tinrhert Plateau in eastern
 366 Algeria and multiple hydraulic jumps on the western side of the TP (locations in Figure 2a). We
 367 discuss the formation of BJ, MGW, and multiple hydraulic jumps in this section. To this end, we
 368 use the vertical cross-sections of the simulated wind, potential temperature, and dust
 369 concentration.

370 4.3.1 Barrier Jet

371 Prior to the formation of the BJ, at 08 UTC on 20 February, the vertical cross-section
 372 along the line A-B shows a near-surface stable layer at $30\text{--}32^{\circ}\text{N}$ (Figure 8a), consistent with the
 373 radiosonde at Béchar (Figure 4a). One can notice the near-surface stable layer with an easterly
 374 wind and vertically veering flow in the observed sounding. The inversion is ~ 350 m deep and the
 375 wind speed exceeds $18 \text{ m}\cdot\text{s}^{-1}$ at some locations already (Figure 8a), which is strong enough for
 376 emitting dust aerosols in source regions. The existence of a low-level stable layer is a critical
 377 prerequisite for BJ formation. An hour later, at 09 UTC, the near-surface wind speed along the
 378 line C-D between 30°N , 3°W and 33°N , 3.3°E largely exceeded $14 \text{ m}\cdot\text{s}^{-1}$ with a maxima $>22 \text{ m}\cdot\text{s}^{-1}$

379 ¹ at $\sim 30.6^\circ\text{N}$, 2°W around 750 m above sea level (ASL) (marked by circle in Figure 8b). This
 380 wind speed maximum is the center of the BJ.

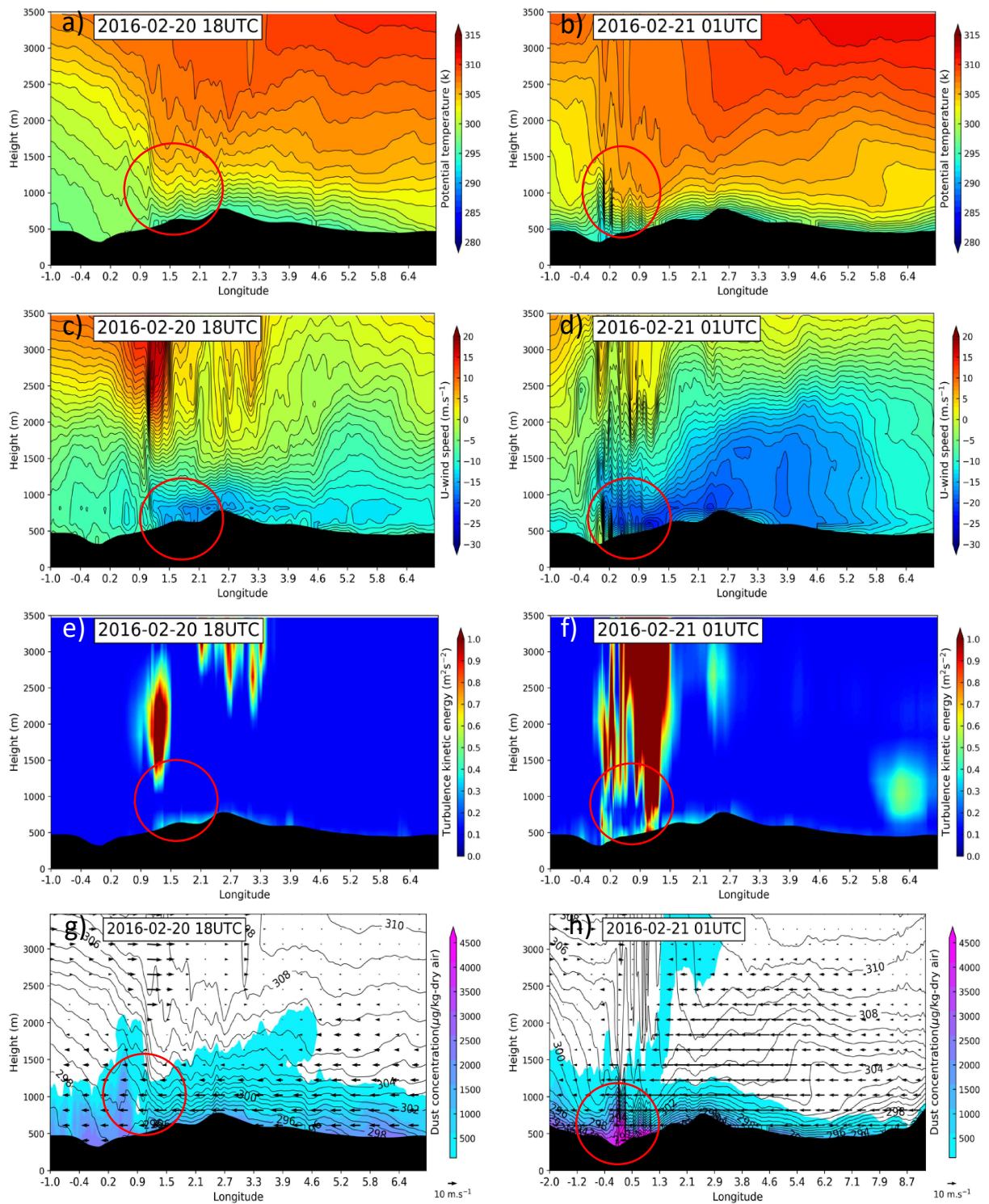


381
 382 Figure 8: Horizontal winds and potential temperature transects for the first dust storm. Shown are
 383 simulation results for the vertical cross-sections of potential temperature (lines in K) and wind
 384 speed (fill in $\text{m}\cdot\text{s}^{-1}$) at (a) 08 UTC along the line A-B and (b) 09 UTC along the line C-D on 20
 385 February 2016. Vertical cross-section of potential temperature and u-wind speed along the line
 386 G-H at (c) 08 UTC and (d) at 09 UTC on 20 February 2016. Location of the transects are marked
 387 in Fig. 5a. The circle in Figure b marks the BJ, in c and d the flow blocking.

388 The typical low-level flow blocking required for BJ formation is seen in the cross-
 389 sections along the line G-H in Figures 8c-d (locations in Figure 5a). At 08 UTC, the time before
 390 a clear signature of the BJ, the low-level flow blocking is seen around 32.38°N , 0.5°W as
 391 indicated by the horizontal gradient in the zonal wind speed (marked by circles in Figures 8c-d).
 392 Here, the low Fr ($=0.29$) indicates insufficient kinetic energy of the air parcel to cross the SAM
 393 steep terrain. At 09 UTC, the Fr was ~ 0.27 at the same location, and the upstream flow remained
 394 in the subcritical flow regime between 08-09 UTC. During this period, the blocked flow
 395 accelerated parallel to the mountain barrier forming the BJ. The BJ formation is further
 396 supported by the highly baroclinic conditions exhibiting strong quasi-geostrophic warm air
 397 advection and veering vertical wind shear (Figures 1b, 1d, 1f, 4a and 5). All these characteristics
 398 are typical for a BJ.

399 4.3.2 Mesoscale gravity wave

400 At ~08 UTC on 20 February, a long-lived westward propagating MGW was triggered
401 near the northern edge of the Tinrhert Plateau in eastern Algeria at ~29°N, 9°E. The MGW
402 reached the western part of the TP by 15 UTC on the same day. The details of the MGW
403 mechanisms are discussed in section 4.4. At 18 UTC on 20 February, when the westward
404 propagating MGW crossed the TP, a signal of a hydraulic jump appeared on its lee side. The
405 sinking and overturning low-level isentropic surfaces at ~1.5-2.1°E mark this hydraulic jump
406 (circle in Figure 9a). The near-surface wind decreased downstream of the jump, while the
407 vertical wind shear is large in the jump region (circle in Figure 9c). The strong change of Fr from
408 0.4 immediately downstream of the jump region at ~1.2°E to 1.67 immediately upstream at
409 ~1.3°E further supports the existence of the hydraulic jump. At 01 UTC on 21 February, there
410 were multiple regions of sub- and supercritical regimes as indicated by sinking and bulging
411 isentropes accompanying multiple differences in wind speed on the western slope of the TP at
412 ~0-1.5°E (circles in Figures 9b and 9d). The situation was further characterized by strong ascent
413 at the leading edge of the jump ~0.2°E indicated by the vertical orientation of the isentropes and
414 large turbulent kinetic energy (TKE) associated with the jump (circles in Figures 9b-f). These
415 features point to the occurrence of multiple hydraulic jumps on the western side of the TP
416 leading to a nocturnal increase in near-surface dust concentrations (Figures 9g-h).



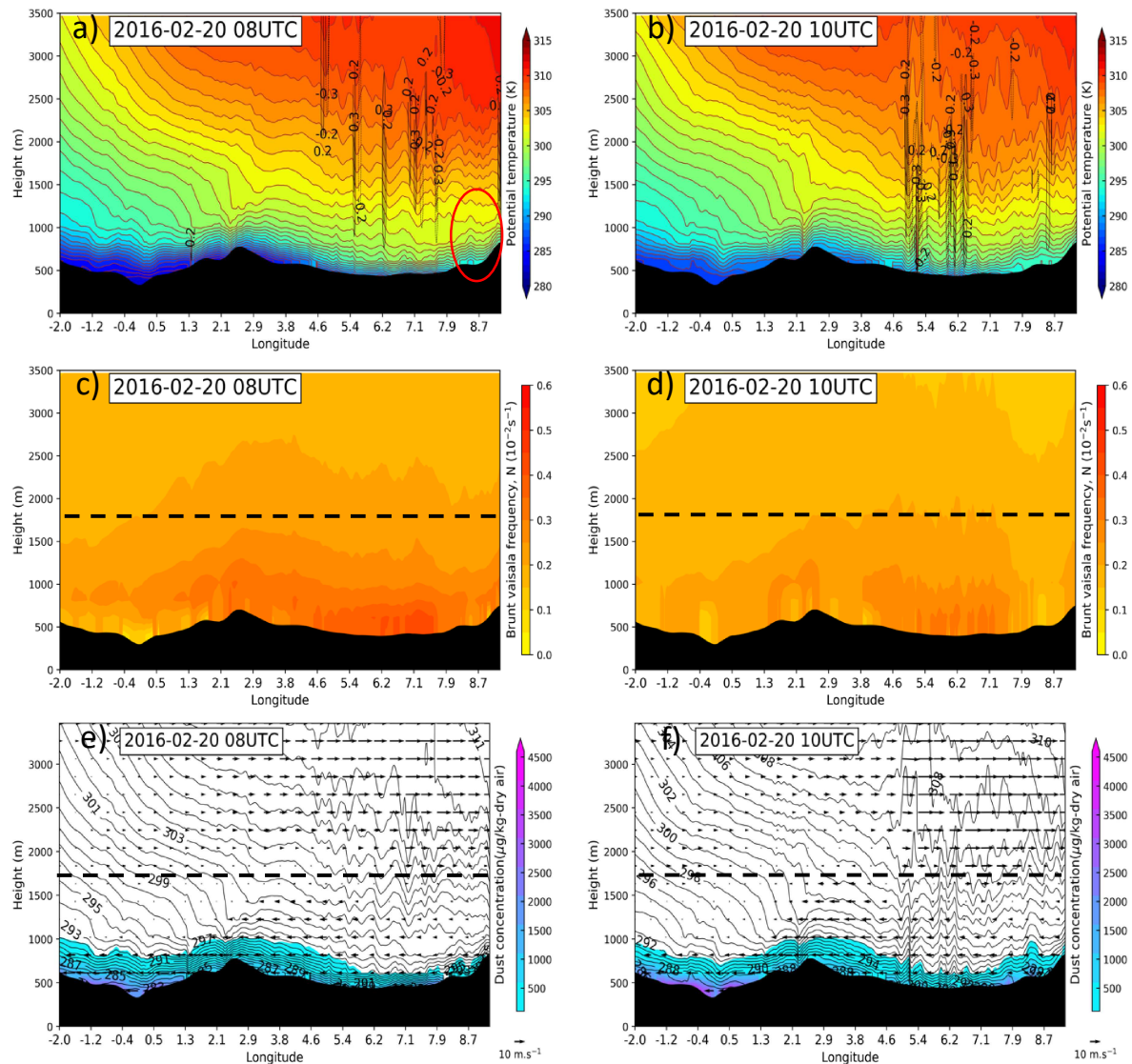
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Figure 9: East-West transect at 29°N for the second dust storm. Shown are WRF-Chem results for the vertical cross-sections at 18 UTC on 20 February (left column) and 01 UTC on 21 February (right column) for (a, b) potential temperature, (c, d) u-wind component, (e, f) PBL TKE (fill in $\text{m}^2.\text{s}^{-2}$), and (g, h) potential temperature (line in K), wind vector (u, w) and dust

422 concentration. The circle marks the region of the hydraulic jump, increasing wind, and dust
423 concentration.

424 4.4 Mesoscale gravity wave generation and maintenance

425 To describe the MGW generation and its maintenance mechanism(s), we analyzed the
426 vertical cross-sections of potential temperature, wind, and Brunt-Väisälä frequency along and
427 extending well beyond the line I-J (Fig. 5a) from 2°W-9°E using the simulation output (Figure
428 10). The MGW was triggered at ~08 UTC on 20 February when cool air descended into the
429 stably stratified air over the lower terrain of the Tinrhert Plateau in eastern Algeria (circle in
430 Figure 10a). The undulating isentropes, together with an ascent-descent pattern at ~5-9°E mostly
431 in quadrature, indicates the existence of a gravity wave (GW, e.g., Uccellini and Koch, 1987).
432 Two hours later, at 10 UTC, the GW propagated continuously westward with a phase speed of
433 ~10.7 m.s⁻¹ (Figure 10b). Atmospheric GWs are highly dispersive in nature and propagate
434 vertically rapidly losing energy which may limit their long-range horizontal propagation from a
435 source region. However, a long-lived MGW like the one that develops here, travels across long
436 distances away from its source region upon satisfying conditions that maintains the wave's
437 energy (Lin, Y., 2007). The wave maintenance conditions in our case follow Lindzen and Tung
438 (1976). They described three key features that are evident for the MGW assessed here as follows:
439 (1) the existence of a near-surface stably stratified layer, indicated by closed isentropes and a
440 high Brunt-Väisälä frequency N (Figures 10a-d), (2) low-level easterly wind, westerly wind
441 aloft, and a mixed layer in between (Figures 10e-f) facilitating a critical layer above the ducting
442 layer near 1800 m, (3) the vertical cross-section of Brunt-Väisälä frequency and wind shear
443 (Figures 10a-d) indicating the low Richardson number $Ri \sim 0.175$ ($N = \sim 0.15 \text{ s}^{-1}$, $dU = \sim 5 \text{ m.s}^{-1}$,
444 and $dz = \sim 1400$) above the duct within the critical layer. The presence of the surface inversion,
445 the critical layer above the duct within the phase reversal velocity at a low and an upper-level,
446 and the $Ri < 0.25$ within that critical layer are consistent with the Lindzen and Tung (1976)
447 criteria for ducting/maintenance. The GW was therefore trapped beneath ~2 km and traveled
448 downwind for a longer duration, i.e., several hundred km and several hours, such that it reached
449 the mesoscale distance and period criterion.



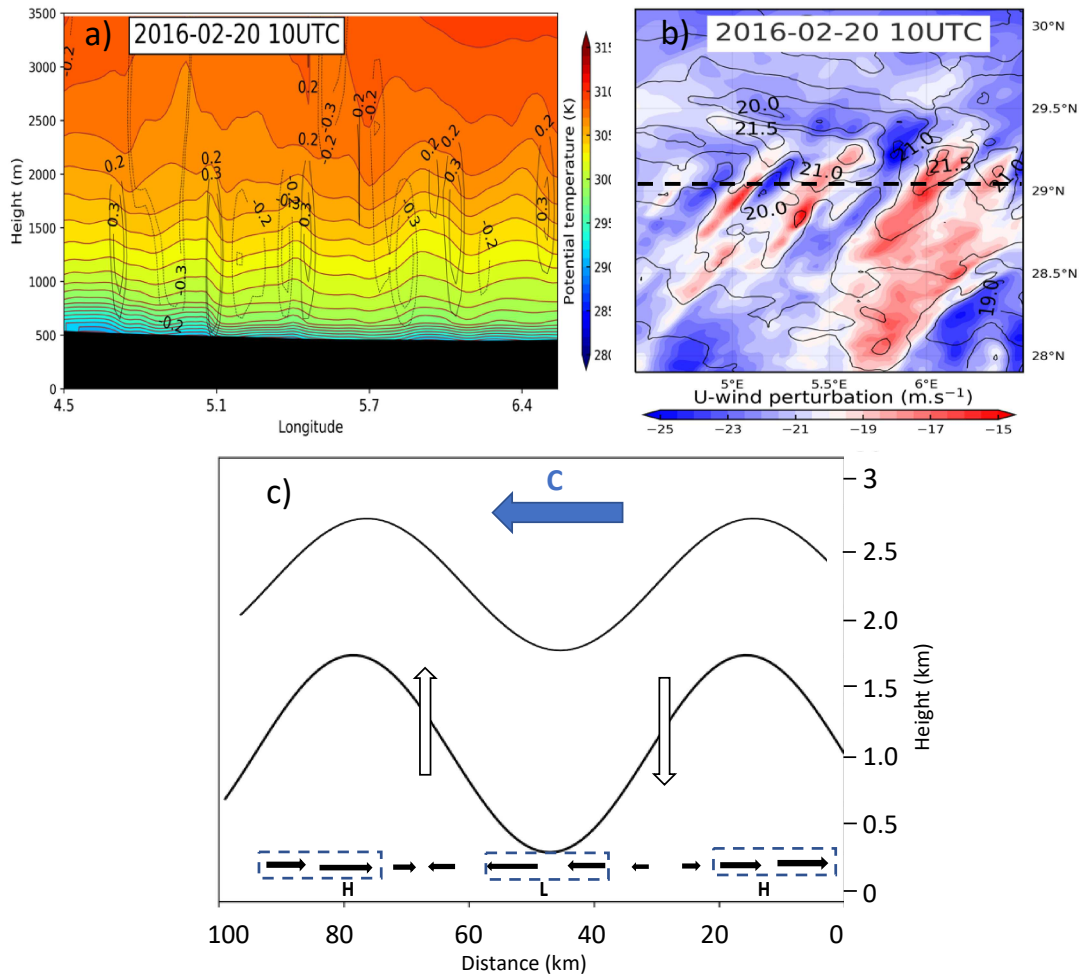
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451 Figure 10: East-West transect at 29°N for the second dust storm. Shown are WRF-Chem results
 452 for the vertical cross-sections at 08 UTC (left column) and 10 UTC (right column) on 20
 453 February 2016 for (a, b) potential temperature and w wind (solid line for positive w and dashed
 454 lined for negative w in $\text{m}\cdot\text{s}^{-1}$), (c, d) Brunt-Väisälä frequency (N , shaded in s^{-1}), (e, f) potential
 455 temperature (line in K), wind vector (u , w) and dust concentration. The dashed line at ~ 1.8 km
 456 represents the wave critical level above the duct.

457

458 Further evidence for the MGW is the relationship between the spatial distribution of the
 459 near-surface pressure perturbation (p') and the wind perturbation (u'). In a ducted gravity wave,
 460 the spatial distribution of p' and u' are positively correlated (e.g., Koch and Golus, 1988;
 461 Coleman and Knupp, 2009). Subsidence warming ahead of a wave trough results in a p'
 462 minimum, while adiabatic cooling results in a p' maximum in the wave ridge (Coleman and
 463 Knupp, 2009). Consistent with the literature, at 10 UTC on 20 February, a positive correlation
 464 between p' and u' was found along the 29°N transect (Figure 11b), illustrating the spatial extent
 of the westward propagating MGW. Noticeable is the evolving high-low alternating pattern of p'

465 and u' , consistent with the pattern of up- and downward motion (Figures 11a-b). The schematic
 466 depiction of the relationship between p' and u' (Figure 11c) indicates that when p' and u' are in
 467 phase, the maximum vertical upward (downward) motion occurs near the nodal point before the
 468 wave ridge (trough). Such a sustained and high amplitude MGW triggers strong near-surface
 469 winds that result in dust emission during its course of motion in a region of several hundred
 470 kilometers.



471 Figure 11: MGW during the second dust storm. Shown are the (a) WRF-Chem simulated vertical
 472 cross-section of potential temperature and vertical wind (solid line for positive w and dashed
 473 lined for negative w in m.s^{-1}) along 29°N at 10 UTC on 20 February 2016, (b) simulated
 474 perturbations in the zonal wind (u' , shading) and pressure (p' , contours in hPa) at the lowest
 475 model level at 10 UTC on 20 February 2016. The dashed black line in b marks the position of the
 476 transect at 29°N shown in a. (c) Schematic diagram to show p' and u' associated with a ducted
 477 gravity wave. The solid black lines represent isentropes, H and L indicate the regions of
 478 maximum and minimum pressure perturbations, horizontal black arrows indicate the perturbation
 479 of the wind vector (maximum and minimum near H and L, respectively), white arrows indicate
 480 the vertical motions, blue arrow indicates the wave propagation direction with the phase velocity
 481 ($C \sim 10.7 \text{ m.s}^{-1}$).
 482

483

484 4.5 Vertical distribution and poleward transport of lifted dust

485 The vertical distribution and poleward transport of dust to the IP were analyzed using the
486 simulated vertical cross-sections of potential temperature, wind, dust concentrations, and 700
487 hPa horizontal winds. At ~09 UTC on 20 February, the emitted dust aerosols from the first dust
488 storm was mostly confined to the near-surface stable layer with a minimal amount reaching 2 km
489 ASL (Figure 7a). The weak southerly wind near ~2 km did not advect a substantial amount of
490 dust poleward. This situation changed with the growing daytime PBL. At 12 UTC, a significant
491 amount of dust was mixed into the deeper convective PBL, reaching altitudes >2.5 km (Figure
492 7b). This dust plume was subsequently advected poleward, which can be proven by the poleward
493 advection of the simulated dust plume and the observed reduction in visibility at the DAAY
494 station (Figures 2f and 7b).

495 The concentrated dust layer from the second dust storm was mostly confined to altitudes
496 <800 m at the foothills of the SAM at 01 UTC on 21 February (Figure 7c). During the following
497 morning, at 06 UTC, the dust layer expanded vertically and reached >2.5 km (Figure 7d). The
498 predominant southerly/southeasterly mid-tropospheric flow advected the available dust poleward
499 and crossed the SAM by midday.

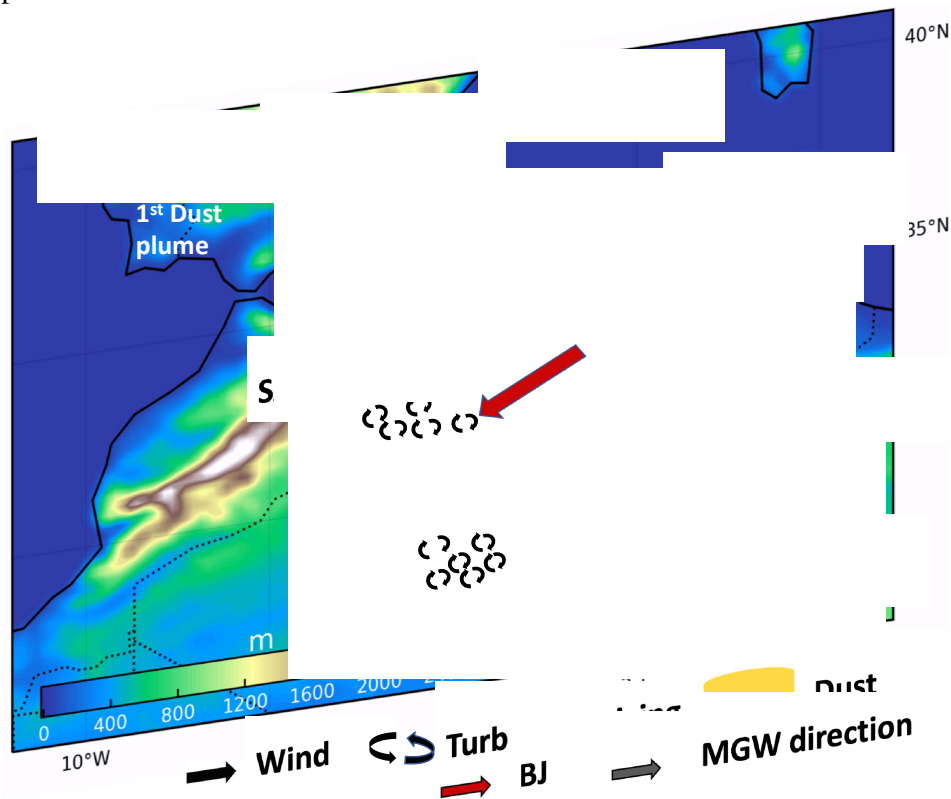
500 5. Summary and conclusions

501 This study investigated the meso- β /meso- γ -scale dynamical processes responsible for the
502 development of the two consecutive severe dust storms over the southeastern part of the SAM
503 during 20-21 February 2016 that impacted the entire IP. The key meteorological findings that
504 organized these dust storms include (1) the evolution of the BJ on the southeastern foothills of
505 the SAM and (2) the development of multiple hydraulic jumps associated with an MGW on the
506 southern flank of the TP. To the best of our knowledge, this study is the first to document the
507 role of a BJ and multiple hydraulic jumps associated with a long-lived MGW in organizing a
508 severe dust outbreak from Northwest Africa towards Europe.

509 Figure 12 shows the schematic depiction of the key mesoscale meteorological dynamics
510 involved in this dust outbreak that had two phases. In Phase I, an easterly/north-northeasterly
511 low-level Mediterranean flow was blocked by the SAM and resulted in BJ on its southeastern
512 foothills. The strong wind associated with the intensification of the BJ after sunrise emitted a
513 massive amount of dust on 20 February 2016 and formed the first dust plume. In Phase II, the
514 MGW affected the western flank of the Tinrhert Plateau in eastern Algeria during the morning of
515 20 February 2016. The MGW triggered multiple hydraulic jumps on the western flank of the TP.
516 The strengthening downslope wind and TKE accompanying the hydraulic jumps enhanced the
517 dust emission on the lee side of the TP and resulted in the second dust plume on 21 February
518 2016. The lifted dust extended over 2-3 km in altitude due to increasing convective turbulence in
519 the growing daytime PBL. The predominantly southerly/southeasterly mid-tropospheric flow
520 subsequently advected the dust plume poleward to the IP.

521 This work underlines the importance of high-resolution numerical modeling in dust storm
522 analyses and operational dust forecasting. The high-resolution model can resolve topographically
523 induced dust-emitting peak winds like the BJ and hydraulic jumps, which are difficult to
524 represent in coarse-resolution (aerosol-) climate and weather prediction models and are rarely
525 observed in the virtually nonexistent meteorological instrumentation in this data sparse
526 unpopulated region. It should be noted here that each dust storm has a unique characteristic

527 concerning the dynamics associated with the dust emission and transport processes. However, to
 528 better quantify the atmospheric dust loading in a longer time span, the impact of climate change
 529 and biological activities over the complex terrain on land cover change need to be addressed
 530 along with other multi-scale dynamical mechanisms critical for dust emission and subsequent
 531 transport processes.



532 Figure 12: Schematic depiction of the time and location of the BJ and hydraulic jumps associated
 533 with MGW and subsequent transport of dust to the IP. Red arrow represents the BJ and gray
 534 arrow represents the MGW propagation direction.

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 544 the ERA-Interim reanalysis dataset, and (4) the University of Wyoming for sounding data.

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